

Ch-3  
Hyper Surfaces

- ①  $\Sigma \in \mathbb{R}^n$       ①  $\phi(x^\alpha) = 0$   
 ② Parametric  
 $x^\alpha = x^\alpha(y^a)$

② Normal  $n_\alpha \propto \partial_\alpha \phi$   
for Non Null  $n_\alpha n^\alpha \equiv \epsilon \equiv \left. \begin{array}{l} 1 \quad \Sigma \text{ spacelike} \\ -1 \quad \Sigma \text{ timelike} \end{array} \right\}$   
Normalized

Convention's  $n^\alpha$  point in the direction of increasing  $\phi$   
 $\therefore n^\alpha \partial_\alpha \phi > 0$  ??  
??  
00

③ From these 2 we can get that

$$n_\alpha = \frac{\epsilon \partial_\alpha \phi}{\sqrt{g^{\mu\nu} \partial_\mu \phi \partial_\nu \phi}} \sqrt{2}$$

④ This will not work for Null case  
 as  $\partial_\mu \phi \partial^\mu \phi = 0$

$$\therefore n_\alpha \longrightarrow \infty$$

⑤ Hence normal can't be Normalized in Null Case

$$\therefore \text{let } k_\alpha = \partial_\alpha \phi$$

when  $\phi \uparrow$  in future  $k_\alpha$  should be future directed

⑥ AS  $k^\alpha k_\alpha = 0 \therefore k^\alpha$  is Tangent to  $\Sigma$

⑦ ~~Null curves are always~~

Th. Given a H.S. Null, we are given Null curves.  
Th. These Null curves always follow Null Geodesic.

Th. H.S. is generated by null geodesics &  $k^\alpha$  is tangent to the generator.

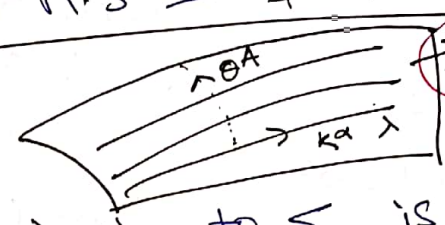
⑧ Th. In general  $\lambda$  is non affine parameter

But when  $\phi = \text{const.}$  describe whole family of null H.S.  
 then  $\lambda = \text{affine.}$

⑨ Coordinates on H.S.  $\equiv$  Intrinsic Coordinates

Null Case

$y^a = (\lambda, \theta^A)$



$\Rightarrow$  2D H.S.  
 $\Rightarrow$  By const.  $k_\alpha e^{\alpha A} = 0$

⑩ The metric intrinsic to  $\Sigma$  is obtained by restricting line element to  $\Sigma$ .

Def:  $e^\alpha_a \equiv \frac{\partial x^\alpha}{\partial y^a} \Rightarrow$  Tangent to  $\Sigma$   
 $\therefore e^\alpha_a n_\alpha = 0$

Why?

⑪ Null case

$e^\alpha_a \equiv \frac{\partial x^\alpha}{\partial y^a} \Rightarrow$  Tangent to  $\Sigma$

$e^\alpha_a = \left( \frac{\partial x^\alpha}{\partial y^1}, \frac{\partial x^\alpha}{\partial y^2}, \frac{\partial x^\alpha}{\partial y^3} \right) = \left( \frac{\partial x^\alpha}{\partial \lambda}, \frac{\partial x^\alpha}{\partial y^A} \right)$

$e^\alpha_a = (k^\alpha, e^\alpha_A)$

$k_\alpha e^{\alpha A} = \text{const}$   
 ut  $k_\alpha e^{\alpha A} = 0$

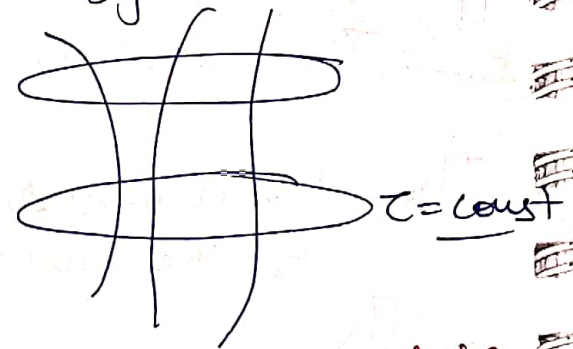


⑫ Non Null

$$dx^\alpha = \frac{\partial x^\alpha}{\partial y^a} dy^a + \frac{\partial x^\alpha}{\partial \tau} d\tau$$

$$ds_\Sigma^2 = g_{\alpha\beta} dx^\alpha dx^\beta$$

$$dx^\alpha_\Sigma = \left( \frac{\partial x^\alpha}{\partial y^a} \right)_\tau dy^a$$



$$ds_\Sigma^2 = (g_{\alpha\beta} e_a^\alpha e_b^\beta) dy^a dy^b$$

$$\equiv h_{ab} dy^a dy^b \Rightarrow$$

$h_{ab}$  is symmetric

$h_{ab} \equiv g_{\alpha\beta} e_a^\alpha e_b^\beta =$  induced metric

⑬ Th.  $\rightarrow h_{ab}$  is scalar w.r.t. transf. of spacetime coord.  
 $\rightarrow h_{ab}$  is tensor w.r.t. transf. of H.S. coord.

$h_{ab} \equiv$  3-Tensor.

we can choose any coord. But lets take this

⑭ Null case

H.S.  $y^a = (\lambda, A) = 3D$

$$e_1^\alpha = \frac{\partial x^\alpha}{\partial \lambda} = k^\alpha$$

$$h_{ab} \equiv g_{\alpha\beta} e_a^\alpha e_b^\beta$$

$$h_{11} = g_{\alpha\beta} k^\alpha k^\beta = 0$$

$$h_{1A} = g_{\alpha\beta} k^\alpha e_A^\beta = k^\alpha e_{\alpha A} = 0 \text{ (by construction)}$$

$$h_{ab} = \begin{pmatrix} 0 & 0 & 0 \\ 0 & h_{22} & h_{23} \\ 0 & h_{32} & h_{33} \end{pmatrix} = \text{Symmetric}$$

$\therefore \underline{h_{23} = h_{32}}$

$h = |h_{ab}| = 0$  as  $h_{ab}$  is Degenerate.

$$\therefore ds_\Sigma^2 = \sigma_{AB} d\alpha^A d\alpha^B ; \sigma_{AB} \equiv g_{\alpha\beta} e_A^\alpha e_B^\beta$$

$\sigma_{AB} \equiv 2$ -Tensor

$e^{\alpha}_A = \left( \frac{\partial x^{\alpha}}{\partial \theta^A} \right)_A$

(15) As  $h^{\alpha\beta} = h^{ab} e^{\alpha}_a e^{\beta}_b$   
 $h^{\alpha\beta} = g^{\alpha\beta} - v^{\alpha} v^{\beta}$   
 $g^{\alpha\beta} = h^{ab} e^{\alpha}_a e^{\beta}_b + \epsilon h^{\alpha\beta}$

Non Null Case

(16) Null Case

why

Introduce  $N^{\alpha}$  everywhere on  $\Sigma$ ,  $N^{\alpha}$  (auxillary Null vectors)

satisfying  $N^{\alpha} k_{\alpha} = +1$

$N_{\alpha} e^{\alpha}_A = 0$

$h^{\alpha\beta} = \sigma^{ab} e^{\alpha}_a e^{\beta}_b$

$g^{\alpha\beta} = \sigma^{AB} e^{\alpha}_A e^{\beta}_B + k^{\alpha} N^{\beta} + N^{\alpha} k^{\beta}$

(17)  $N^{\alpha} N_{\alpha} = 0$   
 $N^{\alpha} k_{\alpha} = 1$   
 $N^{\alpha} e_{\alpha A} = 0$   
 $N^{\alpha} e_{\alpha B} = 0$

4 Cond<sup>n</sup> & 4 variables of  $N^{\alpha}$   
 $\therefore$  solvable

(18) As By construction we can say  $k_{\alpha} e^{\alpha}_A = 0$

let  $k_{\alpha} = (1, -1, 0, 0)$   
 $e_{\alpha 1} = (0, 0, -1, 0)$   
 $e_{\alpha 2} = (0, 0, 0, -1)$

H.S. Basis  
 linear ind.  
Span

$\therefore k_{\alpha} e^{\alpha}_A = 0$

Now let  $k_{\alpha} N^{\alpha} = 1$   $N^{\alpha} = (1, 1, 0, 0)$



$\therefore k_\alpha, N_\alpha$

$a(1, -1, 0, 0) + b(1, 1, 0, 0) =$  They span but they are not linear ind.

(19) As let  $k^\alpha = (1, 1, 0, 0)$   
↓  
for Transverse Direction  
Transv. subspace 2D.  
 $k^\alpha e_{\alpha A} = 0$

(20) Integration on Hypersurfaces

Non Null Case

$h_{ab} = \partial_a y^{a'} \partial_b y^{b'} h_{a'b'}$

$h = J^2 h'$

$d^3 y' = J d^3 y$

$d^2 y' = \sqrt{\frac{h}{h'}} d^2 y$

$\sqrt{h'} d^3 y' = \sqrt{h} d^3 y$

$\therefore \sqrt{h} d^3 y = d\Sigma$  is the surface element.

Directed surface element:  $\eta_\mu d\Sigma$

where  $\eta_\mu \eta^\mu = \epsilon$   
Normalized



(21) Null Case

$$h_{ab} = \begin{pmatrix} 0 & 0 & 0 \\ 0 & \sigma_{ab} \end{pmatrix}$$

$h = 0 \therefore h_{ab}$  Degenerate.

$n_\mu$  is not Normalized

$$n_\mu \rightarrow \infty$$

(22) Therefore till now, we know  
in Non null case  $d\Sigma_\mu = n_\mu d\Sigma = n_\mu \int d^3y$

in Null Case we do not know yet

(23)  $\therefore d\Sigma_\mu = n_\mu d\Sigma$  has to be generalized.  
to take Null Case.

Let  $d\Sigma_\mu = \epsilon_{\mu\alpha\beta\gamma} e_1^\alpha e_2^\beta e_3^\gamma d^3y$   
 $\epsilon_{\mu\alpha\beta\gamma} = \sqrt{-g} [\mu\alpha\beta\gamma]$

(24) Non Null case  
 $d\Sigma_\mu = \epsilon_{\mu\alpha\beta\gamma} e_1^\alpha e_2^\beta e_3^\gamma d^3y \longrightarrow d\Sigma_\mu = n_\mu d\Sigma$

Proof: Claim:  $d\Sigma_\mu = f n_\mu$   
Proof:  $d\Sigma_\mu e_1^\mu = d\Sigma_\mu e_2^\mu = d\Sigma_\mu e_3^\mu = 0$

$\therefore d\Sigma_\mu = f n_\mu$

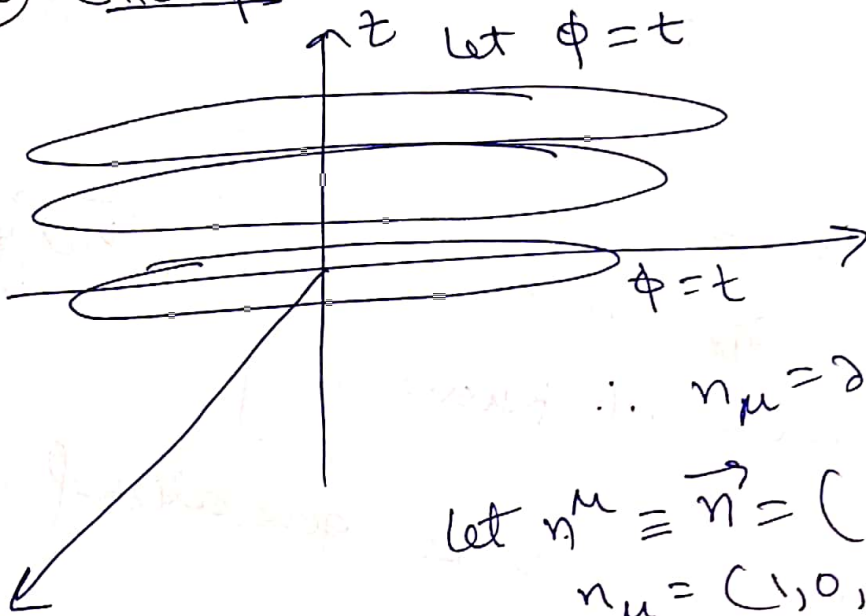
(25) Convention:  $\epsilon_{\mu\alpha\beta\gamma} n^\mu e_1^\alpha e_2^\beta e_3^\gamma > 0$   
 $= \epsilon_{\mu\alpha\beta\gamma} n^\mu \frac{\partial x^\alpha}{\partial y^1} \frac{\partial x^\beta}{\partial y^2} \frac{\partial x^\gamma}{\partial y^3}$

$\epsilon_{\mu\alpha\beta\gamma} n^\mu \frac{\partial x^\alpha}{\partial y^1} \frac{\partial x^\beta}{\partial y^2} \frac{\partial x^\gamma}{\partial y^3}$  depends on the ordering of the 122

Coordinates  $y^1, y^2, y^3$

Convention: Take orderings s.t.  $\epsilon_{\mu\alpha\beta\gamma} n^\mu e_1^\alpha e_2^\beta e_3^\gamma > 0$

② Example: Minkowski spacetime  $\Rightarrow \epsilon_{\mu\alpha\beta\gamma} = [\mu\alpha\beta\gamma]$   
 $\sqrt{g} = 1$



$\therefore n_\mu = \partial_\mu \phi$  (convention:  $\phi$  ↑  $n_\mu$  in that direction)

$$\text{let } n^\mu \equiv \vec{n} = (1, 0, 0, 0)$$

$$n_\mu = (1, 0, 0, 0)$$

Convention:  $\epsilon_{\mu\alpha\beta\gamma} n^\mu e_1^\alpha e_2^\beta e_3^\gamma > 0$

$$[\mu\alpha\beta\gamma] n^\mu e_1^\alpha e_2^\beta e_3^\gamma$$

let  $y^i = (x, y, z)$

$$[\mu 123] n^\mu > 0$$

$$d\Sigma_\mu = \epsilon_{\mu\alpha\beta\gamma} e_1^\alpha e_2^\beta e_3^\gamma dy = f n_\mu$$

$$f = \epsilon n^\mu d\Sigma_\mu$$

$$= \epsilon \frac{\epsilon_{\mu\alpha\beta\gamma} n^\mu e_1^\alpha e_2^\beta e_3^\gamma d^3y}{1}$$

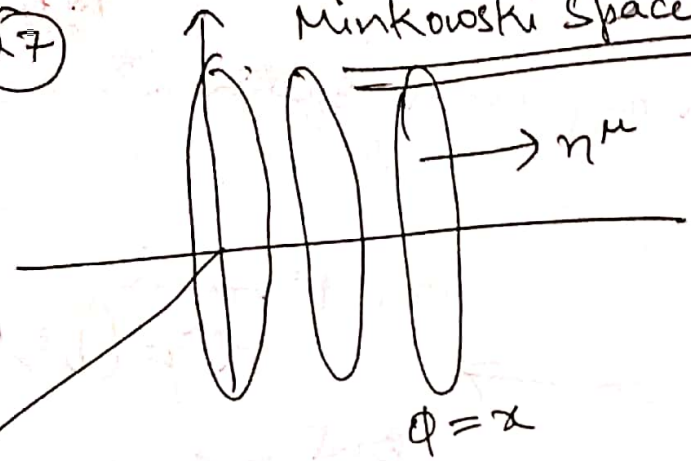
$$f = \epsilon d^3y = dx dy dz$$

$$\boxed{d\Sigma_\mu = n_\mu dx dy dz}$$



Minkowski Spacetime

(27)



$n^\mu = \partial^\mu \phi$   
 But  $n^\mu$  is purely spatial

$\therefore n_\mu = -\partial_\mu \phi$

(convention  $n_\mu$  in direction of  $\uparrow \phi$ )

Let  $n_\mu = (0, -1, 0, 0)$

Convention:

$\epsilon_{\mu\alpha\beta\gamma} n^\mu e_1^\alpha e_2^\beta e_3^\gamma > 0$

$[\mu\alpha\beta\gamma] n^\mu e_1^\alpha e_2^\beta e_3^\gamma$

$[\mu 23 0] n^\mu = \underline{\underline{1}} > 0$

Let  $y^i = (y, z, t)$   
 $\Rightarrow \frac{\partial x^\alpha}{\partial y^i} = \frac{\partial x^\alpha}{\partial y} = (0, 0, 1, 0)$

$d\Sigma_\mu = f n_\mu$

$f = \epsilon d^3y = -dy dz dt$

$d\Sigma_\mu = -n_\mu dy dz dt$

But  $d\Sigma_\mu = n_\mu dy^3$   
 so Be Cautious here

for Timelike H.S.  $d\Sigma_\mu$  in opposite direction of  $n_\mu$

Proof: Non Null Case  $d\Sigma_\mu = \epsilon_{\mu\alpha\beta\gamma} e_1^\alpha e_2^\beta e_3^\gamma dy^3$   
 $\downarrow$   
 $d\Sigma_\mu = n_\mu dy^3$

(28) As we know

$d\Sigma_\mu = \epsilon_{\mu\alpha\beta\gamma} e_1^\alpha e_2^\beta e_3^\gamma dy^3$

$e_i^\mu d\Sigma_\mu = 0 \quad \forall i = 1, 2, 3$

$\therefore d\Sigma_\mu = f n_\mu$

And Also

$\epsilon_{\mu\alpha\beta\gamma} e_1^\alpha e_2^\beta e_3^\gamma = g n_\mu$   
 $g = \epsilon \epsilon_{\mu\alpha\beta\gamma} e_1^\alpha e_2^\beta e_3^\gamma$

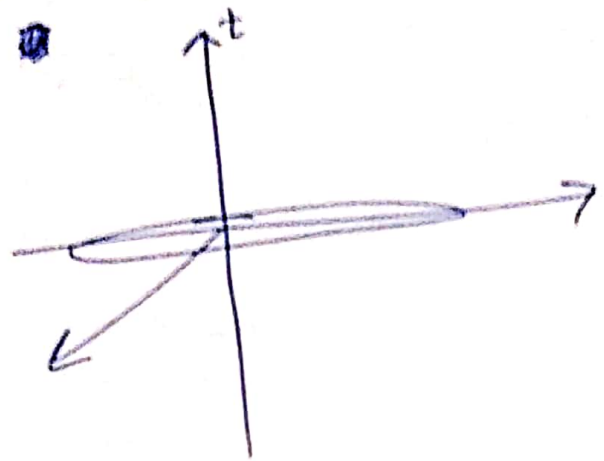
①



As  $g$  is scalar

$\therefore$  we can evaluate in any coord. system  $x^\alpha$ .

let  $\phi = x^0$   
 $y^i = x^i$



Normal  $n^\mu \neq (1, 0, 0, 0)$   
~~But  $\partial_\alpha \phi$~~

$\therefore g \stackrel{*}{=} \epsilon_{\mu\alpha\beta\gamma} n^\mu e_1^\alpha e_2^\beta e_3^\gamma$   
let  $y^i = (x, y, z)$  for  $\epsilon_{\mu\alpha\beta\gamma} n^\mu e_1^\alpha e_2^\beta e_3^\gamma > 0$

$g \stackrel{*}{=} (+1) \sqrt{-g} [\mu 123] n^\mu$

$g \stackrel{*}{=} \sqrt{-g} n^0$

$g^{00} \stackrel{*}{=} \partial_\alpha \phi \partial^\alpha \phi \stackrel{*}{=} g^{\alpha\beta} \partial_\alpha \phi \partial_\beta \phi$  ??

$n_\mu = \frac{\epsilon \partial_\mu \phi}{|\partial_\alpha \phi \partial^\alpha \phi|^{1/2}}$

$n_0 = \epsilon |\partial_\alpha \phi \partial^\alpha \phi|^{1/2} = \epsilon |g^{00}|^{1/2}$

$n^0 = g^{0\alpha} n_\alpha \stackrel{*}{=} g^{00} n_0 \stackrel{*}{=} g^{00} \frac{\epsilon}{|g^{00}|^{1/2}}$

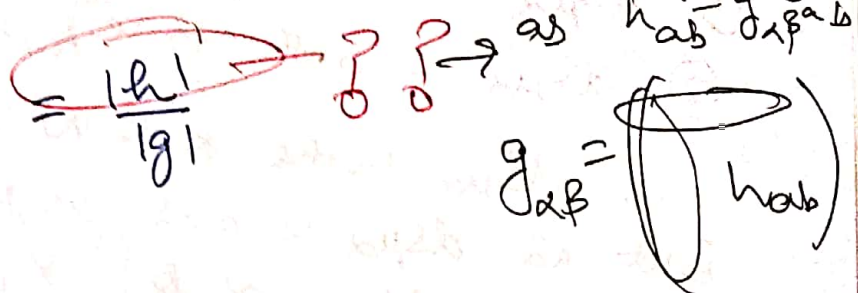
$g^{00} > 0$   
 $\Rightarrow$   
 $\partial_\alpha \phi \partial^\alpha \phi > 0$

AS  $n_\alpha$  is time like  
 $\partial_t \phi > 0$   
 $\partial_\alpha \phi = (1, 0, 0, 0)$   
 $\partial_a \phi = (1, 0, 0, 0)$

$$\therefore n^0 \approx (g_{00})^{1/2}$$

$$g \approx \sqrt{-g} (g_{00})^{1/2}$$

$$g_{00} = \frac{cf \ g_{00}}{|g|} = \frac{|h|}{|g|}$$



$$A^{-1} = \frac{adj \ A}{|A|}$$

$$\therefore g \approx \sqrt{-g} \frac{|h|^{1/2}}{|g|^{1/2}} \approx |h|^{1/2}$$

$$\Rightarrow \text{from ①} \quad \epsilon_{\mu\alpha\beta\gamma} e^{\alpha}_1 e^{\beta}_2 e^{\gamma}_3 = |h|^{1/2} n_{\mu}$$

$$\therefore d\Sigma_{\mu} = |h|^{1/2} n_{\mu} d^3y$$

② Null Case

Choosing our Usual coordinates on null surface  $y^i = (\lambda, \theta^A)$

$$e^{\alpha}_1 = k^{\alpha} \quad e^{\alpha}_{2,3} = \begin{pmatrix} \partial x^{\alpha} \\ \partial \theta_{A,B} \end{pmatrix}$$

$$d^3y = d\lambda d^2\theta$$

$$d\Sigma_{\mu} = \epsilon_{\mu\alpha\beta\gamma} k^{\alpha} e^{\beta}_2 e^{\gamma}_3 d\lambda d^2\theta \quad (\text{Valid for any coordinate system.})$$

$$= k^{\alpha} dS_{\mu\alpha} d\lambda$$

$$dS_{\mu\alpha} = \epsilon_{\mu\alpha\beta\gamma} e^{\beta}_2 e^{\gamma}_3 d^2\theta$$

$dS_{\mu\alpha} \equiv$  2 Dim Surface Area

$$d\Sigma_{\mu} = \epsilon_{\mu\alpha\beta\gamma} e^{\alpha}_1 e^{\beta}_2 e^{\gamma}_3 d^3y$$

Valid for any coordinate system & Any T/S/N case



30 Properties of  $dS_{\mu\alpha}$

1)  $dS_{\mu\alpha}$  is antisym. in  $\mu$  &  $\alpha$ .

2)  $dS_{\mu\alpha}$  is orth. to  $e_A^\alpha, e_B^\alpha$

$\therefore dS_{\mu\alpha}$  in the  $K$  &  $N$

But as  $dS_{\mu\alpha}$  is ant. in  $\mu\alpha$

$\therefore dS_{\mu\alpha} \propto [K_{\mu N_\alpha} - K_{\alpha N_\mu}]$

let  $\epsilon_{\mu\alpha\beta\gamma} e_2^\beta e_3^\gamma = 2f K_{[\mu N_\alpha]}$

31) Moreover, 2D submanifold in 4D needs 2 Normals to be defined

See L-10

26, 27

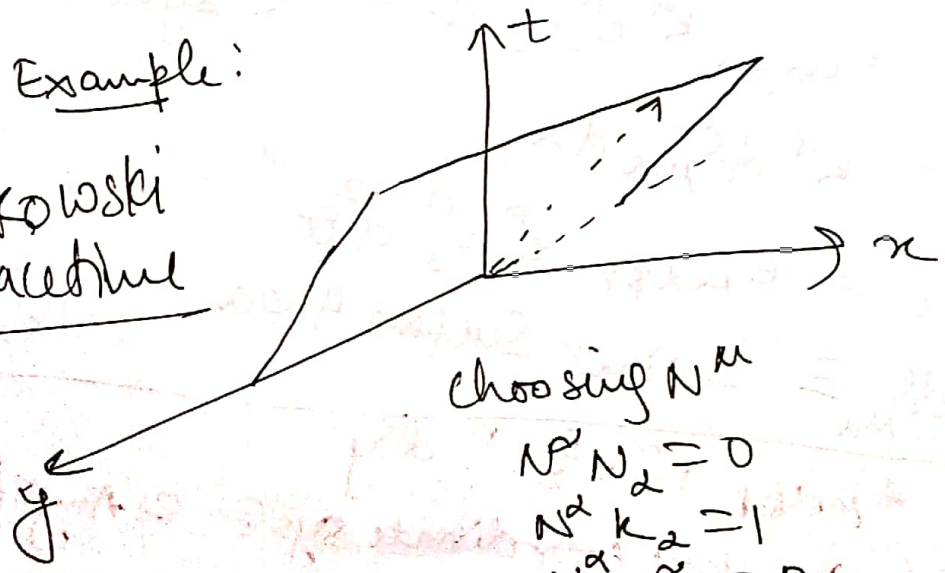
32) What about 1D submanifold in 3D?  
2D submanifold in 3D?

33) Convention:  $\epsilon_{\mu\alpha\beta\gamma} N^\mu K^\alpha e_2^\beta e_3^\gamma > 0$   
i.e. the ordering of  $\Theta^A$  should be s.t.

$\epsilon_{\mu\alpha\beta\gamma} N^\mu K^\alpha \frac{\partial x^\beta}{\partial y^2} \frac{\partial x^\gamma}{\partial y^3} > 0$

34) Example:

Minkowski Spacetime



Choosing  $N^\mu$

$N^\mu N_\mu = 0$

$N^\alpha K_\alpha = 1$

$N^\alpha e_A^\alpha = 0$

$\phi = t - x = \text{const}$

$K_\alpha = \partial_\alpha \phi$

$K_\alpha = (1, -1, 0, 0)$

$e^\alpha = (1, 1, 0, 0)$

$$\begin{aligned}
 \text{let } N_\alpha &= \frac{\partial_\alpha (t+x)}{2} \\
 &= \left(\frac{1}{2}, \frac{1}{2}, 0, 0\right) \\
 N^\alpha &= \left(\frac{1}{2}, -\frac{1}{2}, 0, 0\right)
 \end{aligned}$$

$$\begin{aligned}
 \epsilon_{\mu\alpha\beta\gamma} N^\mu N^\alpha e_2^\beta e_3^\gamma &> 0 \\
 &= [\mu\alpha\beta\gamma] N^\mu N^\alpha e_2^\beta e_3^\gamma \\
 &= \frac{[0\alpha\beta\gamma]}{2} k^\alpha e_2^\beta e_3^\gamma + \frac{[1\alpha\beta\gamma]}{2} k^\alpha e_2^\beta e_3^\gamma \\
 &= \frac{[01\beta\gamma]}{2} e_2^\beta e_3^\gamma - \frac{[10\beta\gamma]}{2} e_2^\beta e_3^\gamma
 \end{aligned}$$

$$\begin{aligned}
 \text{if } \theta^A &= (y, z) \\
 &= \frac{[0123]}{2} - \frac{[1023]}{2} \\
 &= \frac{1}{2} + \frac{1}{2} = 1 > 0
 \end{aligned}$$

$$\begin{aligned}
 \Rightarrow dS_{tx} &= dy dz = -dS_{xt} \\
 d\Sigma_t &= dt dy dz = -d\Sigma_x
 \end{aligned}$$

$$\begin{aligned}
 \sigma_{AB} &= \partial_{A'} \theta_A \partial_{B'} \theta_B \\
 \sigma^1 &= J^2 \sigma \\
 J &= \sqrt{\frac{|\sigma|}{\sigma^1}} \\
 \therefore \sqrt{\sigma} d^2\theta
 \end{aligned}$$

35) As  $dS_{\mu\alpha} = \epsilon_{\mu\alpha\beta\gamma} e_2^\beta e_3^\gamma d^2\theta$

We will show

$$dS_{\mu\alpha} = \left[ 2 k_{[\mu} \overset{\uparrow}{N}_{\alpha]} \right] \left( \sqrt{\sigma} d^2\theta \right)$$

$N_\alpha$  is the auxiliary null vector

gives direction

But we know  $d\Sigma_\mu = k^\beta dS_{\mu\beta} d\lambda$

$$\Rightarrow d\Sigma_\mu = 2 k_\beta k_{[\mu} N_{\beta]} \sqrt{\sigma} d^2\theta d\lambda = k_\mu \sqrt{\sigma} d^2\theta d\lambda$$



∴ for null case

$$d\Sigma_\mu = k_\mu \sqrt{-g} d^2\theta d\tau$$

for non null case

$$d\Sigma_\mu = n_\mu \sqrt{|h|} d^3y$$

$\sqrt{-g} d^2\theta$  : Element of 2 Dim. cross sectional area

$\sqrt{|h|} d^3y$  : Element of 3 Dim cross sectional area

(36)

$$\frac{\partial}{\partial x^\mu} \left( \frac{\partial L}{\partial \dot{x}^\mu} \right) - \frac{\partial L}{\partial x^\mu} = 0$$

$$\frac{\partial L}{\partial x^\mu} = 0$$

$$\frac{\partial L}{\partial \dot{x}^\mu} = 0$$

$$\frac{\partial L}{\partial x^\mu} = 0$$



$$\frac{\partial L}{\partial x^\mu} = 0$$

$$\frac{\partial L}{\partial \dot{x}^\mu} = 0$$

37) ~~As we know~~

Proof:  $dS_{\mu\alpha} = 2 k_{\mu} N_{\alpha} J \sqrt{\sigma} d^2\theta$

As

$$\epsilon_{\mu\alpha\beta\gamma} e_2^{\beta} e_3^{\gamma} = 2f k_{\mu} N_{\alpha}$$

where  $f = \epsilon_{\mu\alpha\beta\gamma} N^{\mu} k^{\alpha} e_2^{\beta} e_2^{\gamma} > 0$

let us choose a coordinate system to evaluate  $f$  as  $f$  is scalar  $\therefore$  we can evaluate it in any system.

let  $x^a = y^a = (\lambda, \theta^A)$  on  $\Sigma$

$$x^0 = t$$

$$k^{\alpha} = \frac{dx^{\alpha}}{d\lambda}$$

$$k^{\alpha} N_{\alpha} = 1 \quad (\text{for choosing 2D submanifold})$$

$$k^{\alpha} = (1, 1, 0, 0) \quad (\text{How?})$$

$$N_0 + N_1 = 1$$

$$N_0^2 - N_1^2 - N_2^2 - N_3^2 = 0$$

(as  $N^{\alpha} N_{\alpha} = 0$ )

but  $N_{\alpha} e^{\alpha}_A = 0$

~~$$N_2 \frac{\partial x^2}{\partial \theta_1} + N_3 \frac{\partial x^3}{\partial \theta_2} = N_2 + N_3 \neq 0$$~~

$$N_0 \frac{\partial x^0}{\partial \theta_A} + N_1 \frac{\partial x^1}{\partial \theta_A} + N_2 \frac{\partial x^2}{\partial \theta_A} + N_3 \frac{\partial x^3}{\partial \theta_A} = N_2 = 0$$

Similarly  $N_3 = 0$

$$\therefore N_0 + N_1 = 1$$

$$N_0^2 - N_1^2 = 0$$

$$N_0 - N_1 = 0$$

$$N_0 = 1/2$$

$$N_1 = 1/2$$

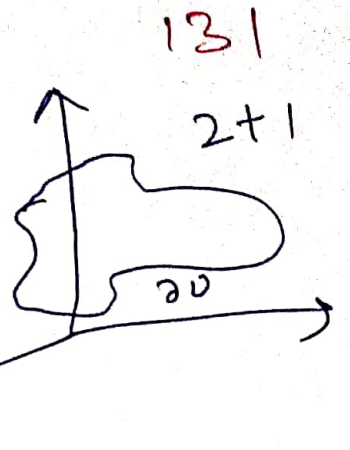
$$\Rightarrow N^{\alpha} = (1/2, -1/2, 0, 0)$$



$$\begin{aligned} \textcircled{37}^0 = K^\alpha K_\alpha &= g_{\alpha\beta} K^\alpha K^\beta = g_{\alpha\beta} \partial^\alpha \phi \partial^\beta \phi \\ &= g^{\alpha\beta} \partial_\alpha \phi \partial_\beta \phi \end{aligned}$$

# 38) Gauss - Stokes Theorem

In Region of  $V$  in manifold  
 Bounded by closed H.S.  $\partial V$   
 The H.S. may have segments  
 that are timelike, spacelike or null



$$\int_V \nabla_\alpha A^\alpha \sqrt{-g} d^4x = \oint_{\partial V} A^\alpha d\Sigma_\alpha$$

where  $d\Sigma_\mu = \epsilon_{\mu\alpha\beta\gamma} e_1^\alpha e_2^\beta e_3^\gamma d^3y$